

## **Physics at CERN** or how to find answers to the fundamental questions of the Universe?

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Since the dawn of time, humans explore the world around and try to explain the origin of the Universe, its evolution and everything it contains.

For a long time, science was mixed up with religious believes, philosophical considerations and factual observations.



## Renaissance: a new era for science

Galileo Galilée – 1564-1642 : « Measure what is measurable, and make measurable what is not so »

# ALLIARI INVENTIVA ET OPVS.OV

#### Antoni van Leeuwenhoek – 1632-1723 :

« My work, which I've done for a long time, was not pursued in order to gain the praise I enjoy now, but chiefly from a craving after knowledge, which I notice resides in me more than in most other me









#### The structure of matter





## From a particle zoo to the Quark Model

**1940-1960:** First hints of  $\pi$ -mesons in cosmic ray picture (Bibha Chowdhuri)

→ discovery of many hadronic particles : π, K, Λ, Δ, Ξ, ρ, φ, Σ, η, ω, Ω...

Enrico Fermi to his student Leon Lederman: "Young man, if I could remember the names of these particles, I would have been a botanist."

→ searching for underlying structure in a zoo of particles: quest for beauty and simplicity?

#### Murry Gell-Mann and Kazuhiko Nishijima (1961) : → "Eight-fold Way" Gell-Mann and George Zweig (1964):

- $\rightarrow$  model with an underlying structure of 3 quarks
- → Considered first as a mathematical construct with no underlying physical reality





## QCD: Quantum Chromodynamics

#### Experiment at SLAC in 1968 (Jerome Friedman, Henry Kendal, Richard Taylor)

→ The elastic cross-section would be expected to drop at higher energy, but it stays stable: it "scales"

#### James Bjorken in 1967:

"scaling" indicates « that the nucleon consists of a certain number of 'elementary constituents' » Friedman's comment:

« Bjorken's results were based on current algebra, which we found as highly esoteric »

#### **1973 : Harald Fritzsch and Heinrich Leutwyler**

proposes that quarks carry a **color charge David Gross, David Politzer and Frank Wilczek** discover **asymptotic freedom** explain why quarks are not observable as free particles but behave as such in high energy interactions



Wilczek : « In theoretical physics, paradoxes are good... When our physical theories lead to paradox we must find a way out. Paradoxes focus our attention, and we think harder »





## Proton structure at HERA (1991-2007)

#### .. Some 40 years later:

Results from Deep Inelastic Scattering: HERA collider

- → Observation of Scaling when the electron interacts with a constituent quark of the proton
- → Observation of scaling violations when the electron interactions with a gluon or a quarks from quark-antiquark production







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## Parton distribution functions



Combination of data from Deep inelastic scattering and other processes observed in different experiments (jet production, single flavour production...)

Parton distribution function of the proton :
→ used in all simulations to describe interactions at LHC
→ used for air show simulation of cosmic rays
Observation : gluon density rises dramatically at very small x
→ Will there be a saturation effect?

 $\rightarrow$  Further improvements from QCD measurements at LHC!

Also f.ex. at CERN: Compass experiment (taking data since 2002):

- Looking at the proton at lower energy: proton tomography
- → Resolving the proton structure in the position space
- Generalized Parton Distribution functions
- → Soft particles surround a hard core of valence quarks



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#### Instruments in HEP

Particle Physics requires huge instruments:

- Particle colliders
- Particle detectors

Every instrument is unique, built by the collaboration in the participating institutes all over the world based on a best effort principle

#### Detection principle: mainly ionisation

- in low density materials → tracking possible for charged particles electrical charge measurement du to magnetic field
- In high density materials  $\rightarrow$  calorimetry

Aim: gather as much information as necessary and as few as possible



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#### Calorimeters



- Design and construction
- Electronics to trigger and read-out the signals
- Calibration system to ensure linearity and uniformity of the signals
- Reconstruction to convert electric signal to energy
- Reduce "noise"
- Identify objects as electrons, photons, jets
- Combine with information from other detectors
- → Major effort to operate the detectors on a daily basis over 2-3 decades



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#### Top quark



Quarks in the Standard model: 6 quarks with masses ranging from ≈ 2MeV – 173 000 MeV (≈3 10<sup>-23</sup> kg) Quark discoveries: c-quark → J/Psi meson 1974 by Burton Richter and Samuel Ting b-quark → Y meson 1977 by Leon Lederman top-quark → 1995 by CDF and D0 collaborations

Top quark properties:

- mass as high as a gold-atom!

Ifetime so short, it cannot create a bound state
 → In the Standard Model, the Higgs mass can be computed from the masse of the W-boson and the top-quark: prediction and consistency check!
 →LHC: top factory!





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> > М<sub>∾</sub> [G

80.5

G fitter

180

190

m, [GeV

170

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≥935

## Higgs Discovery

Announcement on the 4<sup>th</sup> of July 2012 by the Atlas and CMS collaborations : 6000 physicists sign the discovery publications The standard model is completed !





100



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## It's been a long way...

VOLUME 13, NUMBER 16

PHYSICAL REVIEW LETTERS

19 October 1964

#### BROKEN SYMMETRIES AND THE MASSES OF GAUGE BOSONS

Peter W. Higgs Tait Institute of Mathematical Physics, University of Edinburgh, Edinburgh, Scotland (Received 31 August 1964)

about the "vacuum" solution  $\varphi_1(x) = 0$ ,  $\varphi_2(x) = \varphi_0$ : In a recent note<sup>1</sup> it was shown that the Goldstone theorem,2 that Lorentz-covariant field  $\partial^{\mu} \{\partial_{\mu} (\Delta \varphi_1) - e \varphi_0 A_{\mu}\} = 0,$ (2a) theories in which spontaneous breakdown of symmetry under an internal Lie group occurs contain zero-mass particles, fails if and only if  $\{\partial^2 - 4 \varphi_0^2 V''(\varphi_0^2)\}(\Delta \varphi_2) = 0,$ (2b) the conserved currents associated with the internal group are coupled to gauge fields. The  $\partial_{\nu}F^{\mu\nu} = e\varphi_0\{\partial^{\mu}(\Delta\varphi_1) - e\varphi_0A_{\mu}\}.$ (2c) purpose of the present note is to report that, as a consequence of this coupling, the spin-one A PHENOMENOLOGICAL PROFILE OF THE HIGGS BOSON

John ELLIS, Mary K. GAILLARD \* and D.V. NANOPOULOS \*\* CERN, Geneva

Received 7 November 1975

We should perhaps finish with an apology and a caution. We apologize to experimentalists for having no idea what is the mass of the Higgs boson, unlike the case with charm [3,4] and for not being sure of its couplings to other particles, except that they are probably all very small. For these reasons we do not want to encourage big experimental searches for the Higgs boson, but we do feel that people performing experiments vulnerable to the Higgs boson should know how it may turn up.

**1964 : Independent publications of theoretical papers** by Brout and Englert and Higgs on "Electroweak symmetry breaking".

... to disover the Higgs boson

1975: publication by Ellis, Gaillard, Nanopoulos2013: Noble Prize to François Englert and Peter Higgs





## It's been a long way...

ECFA 84/85 CERN 84-10



held at Lausanne and Geneva,

21-27 March 1984

#### ... to built the LHC!

1984 : founding workshop in Lausanne
1996 : LHC approval: total cost about 5€
2008 : first LHC- beam → Incident
2009: first collisions
≈2025: High-luminosity LHC
≈2038: end of data taking





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## Higgs boson: been there, done it?



# Or: why a discovery is just the beginning

- Higgs boson completes the standard model
- Why is it interesting to study the Higgs?
- What can we learn with more statistics?





## Imagine: the Higgs field

The Universe is filled with fields. In the vacuum pairs of particles-antiparticles are constant created and annihilate The mean value of the energy of this process is 0.

The Universe is also filled with the Higgs field. It is a scaler field: no electric charge, no spin. In the vacuum, scaler fields can have a nonzero energy. Higgs field: <vev> = 246 GeV

The Higgs particle is the "quantum excitation" associate to this field. All massive particles interact with the Higgs: the strength of their coupling determines their mass

**BUT:** 99% of the mass of particles such as protons and neutrons is due to the binding energy of QCD

## Higgs: scaler field and inflation?



#### Inflation could describe the very early expansion of the Universe:

- Expansion would be due to the scaler inflaton-field
  - scaler fields have a ground state with an energy ≠ 0 !
  - potential of the inflaton-field would decrease over time
     creation of matter-antimatter particles creates
     negative pression

#### Good news: scaler fields exist !

Bad news: doesn't look to be the Higgs field : Higgs is to heavy Higgs-inflaton oscillations → not observed (yet) at the LHC

#### Future investigations on inflation:

- Inflation should have produced primordial gravitational wave
- Polarised CMB

or observation of cosmic gravitational wave background

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 $10^{-5}$ 

10<sup>-15</sup>

10<sup>-20</sup>

 $10^{-25}$ 

10<sup>-30</sup>

**RADIATION-**

DOMINATED

Dark energy density

10

Time since Big Bang (yr)

Density (kg/m

## Expansion of the Universe



DARK

ENERGY-DOMINATED

Matter-radiation

crossover point

MATTER-

DOMINATED

Matter

density

 $10^{8}$ 

 $10^{10}$ 

- Expansion of the Universe theorized in the beginning of the 20<sup>th</sup> century
   What causes the expansion?
- Vacuum energy? ≈10<sup>120</sup> times to high !!! « worst theoretical prediction ever »
- Accelerated expansion of the Universe discovered in 1998
- $\rightarrow$  Noble Prize for Perlmutter and Riess in 2011

 $\Lambda$ CDM model: accelerated expansion due to "dark energy" – corresponding to a cosmological constant  $\rightarrow$  scaler field?

Expansion of the Universe quantified by the Hubble Constant:Expansion rate measurements show disagreement between"near sources" (Cepheides) and "far away sources" (CMB, BAO, SN)



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#### Higgs boson: mass measurements

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Measurement of the Higgs mass at the 1 ‰ level !
→ Higgs mass measurement: 125.1 +/- 0.14 GeV

Top mass measurement: 172.8 +/- 0.3 GeV Top, W and Higgs mass  $\rightarrow$  consistency check

#### Is the vacuum state of the Universe stable?

Evolution of the vacuum up to the Planck scale:
→ Calculation based on the Higgs mass and the top mass
Current measurement indicate meta stability
→ More accurate measurement (top mass!) → ee-collider
→ New physics?
→ Or Universe may transit to another vacuum state at some point





## Higgs couplings



If the particle discovered at Atlas and CMS is the Higgs boson, the coupling to elementary particles (quarks, leptons, bosons) has to be proportional to their masse !

→ The higher the masse, the easier to verify
 → Verification down to the 2<sup>nd</sup> generation

For physics beyond the Standard Model  $\rightarrow$  deviations < 1%



## Higgs portal to Dark Matter?

Dark Matter first proposed Fritz Zwicky in the 1930
→ firmly established by Vera Rubin in the 1970s
All massive particles should couple to the Higgs boson:
Could we see Dark Matter particles through the Higgs boson?

 1) Higgs mass → life-time of the particle
 → particle width Γ<sub>H</sub> = 4.1MeV
 → related to the couplings Current limits: ≈ 3-4 x SM value

2) Search for invisible particles in Higgs decays





FOR ASTROPHYSICS, ET AL



## Higgs self-coupling





Higgs boson has a mass  $\rightarrow$  it couples to itself!

The Higgs boson in the Standard Model has a potential of:

 $V = -m_{\rm H}^2 |\phi|^2 + \lambda |\phi|^4$   $\rightarrow$  vev = 246 GeV computed from  $M_{\rm W}$  and the weak coupling g  $\rightarrow$  Relation vev,  $m_{\rm H}$  and  $\lambda$  (coupling)  $\lambda = m_{\rm H}^2/2 \text{vev}^2 \approx 0.13 \Rightarrow$  strong prediction

Potential deviations:

- Beyond Standard model physics



## Higgs phase transition



Higgs potential must have evolved during the cooldown of the Universe : How?

2<sup>nd</sup> order phase transition: Universe stays in thermic equilibrium → favoured by the Higgs mass measurement ( $m_H$ <70 GeV): matter creation?

#### 1<sup>st</sup> order phase transition:

No thermic equilibrium  $\rightarrow$  formation of "bubbles" and tunnelling effects : required for baryogenesis, but physics beyond the Standard Model necessary i.e. additional scaler fields



#### Where is the antimatter?

In the current Universe: 10<sup>9</sup> photons, 1 proton - 0 antiprotons Yet, matter and antimatter should have been created in equal amounts!

> 1 μs after the Big Bang: 500 000 001 quarks and 500 000 000 antiquarks annihilate!

> > Why does matter exist ?

- **3** Sakarahov conditions for baryogenesis:
- 1) No thermal equilibrium:

reaction and backwards reaction not in equilibrium

 $\rightarrow$  1<sup>st</sup> order phase transition  $\rightarrow$  measurement of Higgs potential

#### 2) Non-conservation of the baryon number:

stability of the proton measured to  $10^{31}$ : limit to be reached  $10^{39}$   $\rightarrow$  Sphaléron: non-perturbative process at high energy in the Standard Model  $\rightarrow$  transformation of 3 baryons into 3 anti-leptons  $\rightarrow$  could this be observed at a collider?



## CP-violation: Anti-matter in the mirror

#### 3) CP violation:

Charge-Parity symmetry is not conserved : matter and anti-matter behaves differently Indirect CP violation observed by Cronin and Fitch in 1964 in Kaon decays Direct CP violation observed in 1990 at CERN and Fermilab for Kaons in the 2000 by BaBar and Belle for B-mesons, today LHCb, Belle-2 Baryogenesis: CP violation from weak interactions in the quark sector too small x 10<sup>9</sup> Other sources of CP violation? strong CP problem: no CP violation in strong interactions, but theories allows for!

Measurement: electric dipole moment of the neutron → new measurements to come from PSI Solution proposed in 1970s by Peccei and Quinn → a new particle : the Axion



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## Antimatter - antigravity



#### Or... Could anti-matter behave different than matter for gravity?

First ideas in the 1960's form Roland Omnès Today work on a model of "Dirac-Milne Universe" – to be compatible with current observations from cosmology

#### And experiments?

 difficulty: produce "cold" antihydrogen in order to achieve sensitivity for gravitational effects
 → several experiments in preparation at CERN "Antimatter Factory"





## Neutrino physics

Neutrinos predicted by Wolfgang Pauli en 1930 → energy conservation in radioactive beta-decays: *"I have done something very bad today by proposing a particle that cannot be detected; it is something no theorist should ever do."* 

 $\rightarrow$  v discovered in 1956 by Reines and Cowan close to a nuclear reactor ( $v_{\mu}$  in 1962 and  $v_{\tau}$  in 2000)

Solar neutrino problem: flux of solar neutrino flux observed by the Homestake experiment (1968 -2000): shape corresponds to prediction by Behcall but flux only 1/3
 → Discovery in 2001 of neutrino oscillations by SNO and Super-Kamiokande (first ideas on neutrino oscillation: end 1950's by Bruno Pontecorvo)

Implication: only massless, left-handed neutrinos (spin/chirality) exist in the Standard Model!





## Neutrinos : the unknows

Accelerator based, reactor based and atmospheric/cosmic neutrino experiments ongoing!

- Measurement on the oscillation parameters: mass hierarchies, mixing angles
- Search for possible CP violation → would be a hint towards "leptogenesis models"
- Search for possible righthanded neutrinos :
  - Could be sterile neutrinos?
  - Could be very heavy neutrinos (see-saw mechanism) ?
  - If so, dark matter candidates?
- Could the neutrino be its propre anti-particle? Dirac or Majorana type neutrinos?
- Absolute Neutrino mass measurements
  - → Katrin experiment, model-dependent constraints from cosmology
- Dream: Cosmic neutrino background !









#### Enrico Fermi's ultime accélérateur



Vision of Enrico Fermi in 1954 : principle of a particle collider was not know → beams were only shot on targets :

- Accelerator of ≈100 TeV
- Should be built by 1994
- Price: 170 Md\$
- Preliminary concept: 8000 km magnets of 2 T

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## What we think of today...

CERN



## CERN

#### ESPP conclusions



2020 UPDATE OF THE EUROPEAN STRATEGY FOR PARTICLE PHYSICS

by the European Strategy Group

European strategy for particle physics est. 2005:

- updated every 5-7 years by the CERN Council
- bottom-up process with about 160 contributions received and discussed during an Open Symposium
- $\rightarrow$  Produced a 200 page briefing book
- → Transvers working groups: early careers and diversity, global governance, relations with other fields, technology transfer, communication and outreach, environmental impact
- "topish-down": delegates from all European Member States of CERN formulate the strategy

#### Questions from participants

- Comments and discussion on HE flavour opportunities at *pp* vs *e<sup>+</sup>e<sup>\*</sup>*.
   Is it reasonable to have HE flavour effects in view of low-
- energy constraints? 9. Should CERN take any role in being a hub for
- technology/experiment/theory/computing towards DM searches?





## High priority future initiatives

- An electron-positron Higgs factory is the highest priority
- Ambition at the longer term to operate a hadron collider at the highest achievable energies
- Development of innovative accelerator technologies : muon-collider, plasma acceleration, two-beam acceleration

#### What is needed:

Development of high field superconductive magnets, if possible at high critical temperature
 Technical and Financial feasibility study, including environmental impact





ORGANISATION EUROPÉENNE POUR LA RECHERCHE NUCLÉAIRE CERN EUROPEAN ORGANIZATION FOR NUCLEAR RESEARCH

#### CONVENTION

#### FOR THE ESTABLISHMENT OF A EUROPEAN ORGANIZATION FOR NUCLEAR RESEARCH



Cost of a new collider needs additional resources than current CERN budget

- First global infrastructure?
- Contributions from the private sector?
- Scientific research is motivated by scientific questions
  - Innovation is essential but not the first goal
- Particle Physics requires long term

#### engagement

Coordination structures and orgnisations such as CERN are essential

#### Reducing environmental impact

- Primordial, but possible field for innovation



### Quests in particle physics

- Particle physics aims to understand the basic properties of particles, and to describe the evolution of the universe at a fundamental level
- Strong link to cosmology and quantum physics
- Fundamental questions: is the Standard Model valid up to the Planck Scale?
  - Can we built and test a theory of quantum gravity?
  - How to explain matter-anti-matter asymmetry?
  - Does dark energy and dark matter have a microscopic explanation?
- Two fold way: precision measurements and exploration of high energy frontier
- How to go there?
  - $\rightarrow$  Explore the feasibility of colliders we know in principle how to built
  - $\rightarrow$  Explore the possibility of colliders we don't know how to built

# CERN

## What I like about working in particle physics

#### Particle physics is collaborative !

- No one can obtain a result by themselves: thinking of the common good in a collaboration is mandatory
- Contributions of everyone are recognized through authorship
- Possibility to experience a large variety of activities: Hardware, software, analysis, phenomenology, management, communication...
- Flat hierarchies and open to initiatives and ideas, autonomy
- International coordination to avoid duplications
- Promotion of Open Science

